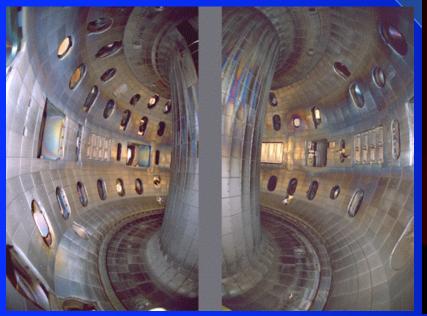
The following slide show is a compilation of slides from many previous similar slide shows that have been produced by different members of the fusion and plasma physics education community. We realize that some of the information contained herein must be updated. Please send comments, complaints, and suggestions to: rick.lee@gat.com. This slide show is intended to be used by students and teachers; downloading this file for educational

purposes is highly encouraged.

Plasma: the 4th State of Matter and a Path to Fusion Energy

What is a plasma? Why should we care? How can we make fusion work? Where are the difficulties?







- What is a plasma?
- Where do we find them?
- Why are we interested in them?
 - Fusion energy
 - Astrophysics
 - Plasma processing

What is a plasma?

- A plasma is an ionized gas.
- Plasma is called the "fourth state of matter."
- Much of the mass of the universe is in the plasma state.
- 'Plasma' was coined by Tonks and Langmuir in (1929):

"...when the electrons oscillate, the positive ions behave like a rigid jelly..."

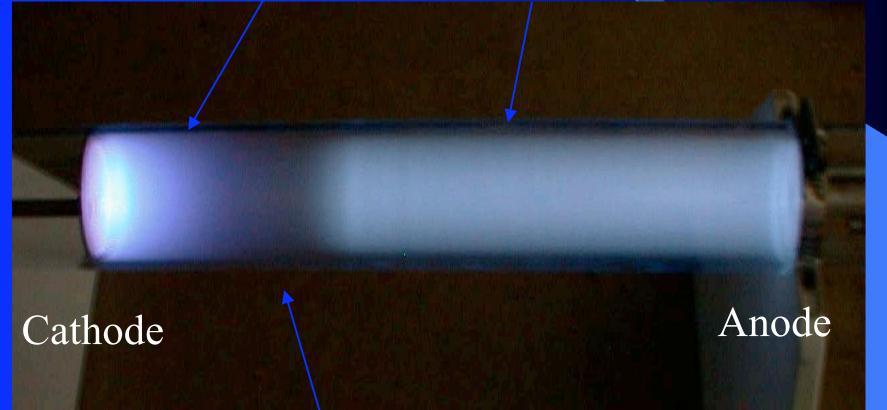
Commercial plasma balls contain mixtures of helium, neon, argon, and other gases. Once the power is switched on, some of the neutral gas molecules have one or more electrons ripped from them, thus producing

Plasma.....



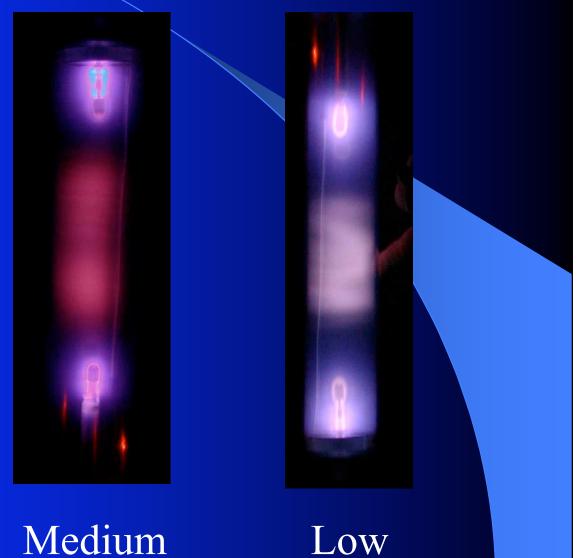
Classical DC Discharge in a Tube

Positive Column



Faraday Dark Space

Plasma is Affected by Density and Mean Free Path



High

Medium

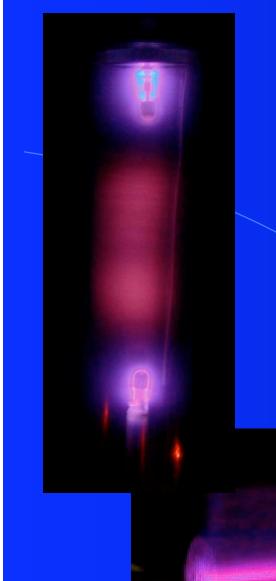




- Same as large bulb except
 - Current density is higher
 - Sound is louder---what does this mean?
 - Colors are similar to the large bulb but more intense

<u>Drift</u>

- The motion of charge particles
 - Random thermal activity
 - Electric field forces
 - Mobility decreases as pressure increases
 - After 1 electron mean free path, an electron may produce one electron-ion pair
 - Ion pairs grow exponentially with electron distance

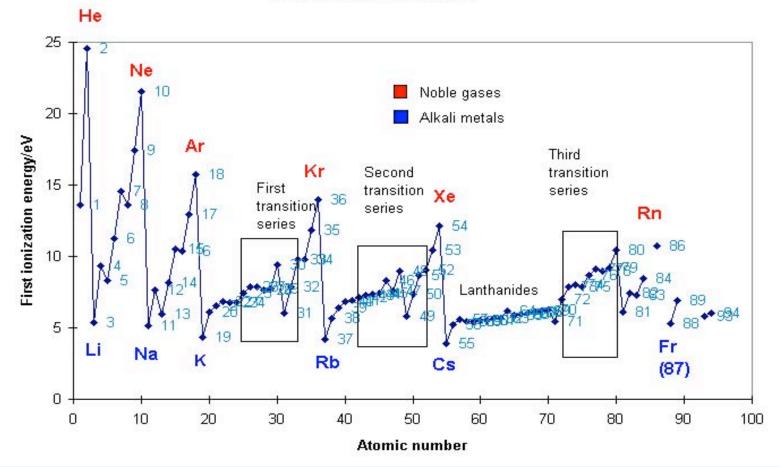


AC Plasma •Typically used for lighting •Both ends become the cathode and anode

•At high pressures the dark spaces are thin and current densities high

Ionization Energies

First Ionization Energies



Where do we find plasmas?

Examples of plasmas on Earth:
Lightning
Neon and Fluorescent Lights
Laboratory Experiments
Examples of astrophysical plasmas:
The sun and the solar wind
Stars, interstellar medium

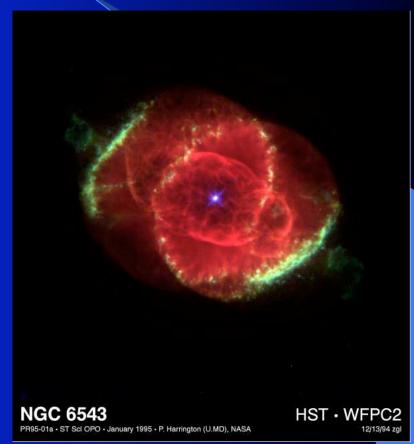
Astrophysical plasmas

Catseye Nebula

The Sun



http://bang.lanl.gov/solarsys/



http://www.stsci.edu:80/

Plasmas on Earth

Laboratory Experiments



http://FusEdWeb.pppl.gov/

Lightning



Why are we interested in plasmas?

• Fusion Energy

- Potential source of safe, abundant energy.

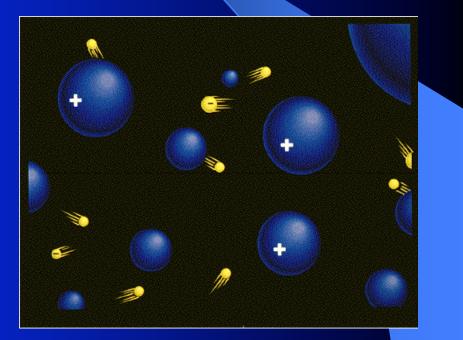
• <u>Astrophysics</u>

- Understanding plasmas helps us understand stars and stellar evolution.
- Upper atmospheric dynamics
 - The upper atmosphere is a plasma.
- Plasma Applications
 - Plasmas can be used to build computer chips and to clean up toxic waste.

Properties of plasmas

- A collection of positively and negatively charged particles.
- Plasmas interact strongly with electric and magnetic fields.
- Plasmas support many different types of waves and oscillations.

Cartoon of a plasma

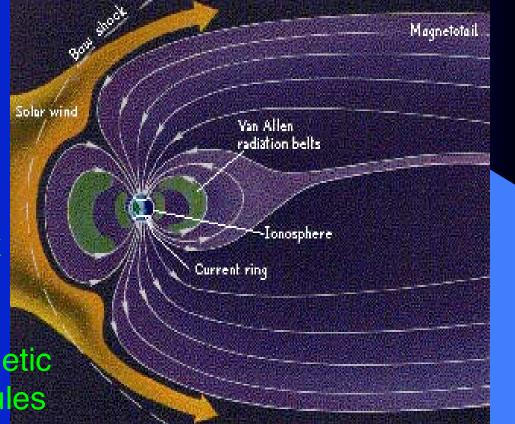


The solar wind (a plasma) interacts with the Earth's magnetic field

The sun emits mass in the form of plasma at velocities of up to 500 km/s.

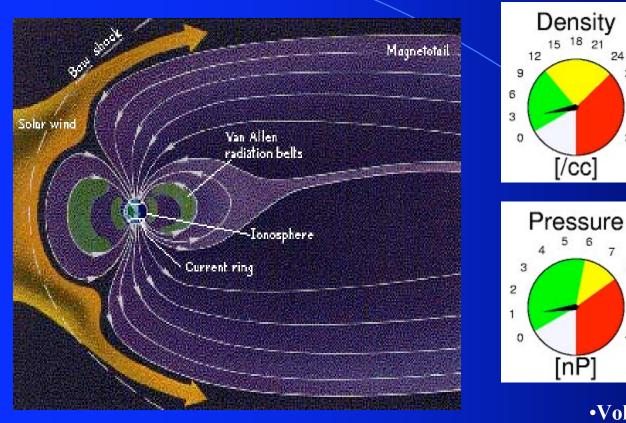
This solar wind causes the Earth's magnetic field to compress creating a shock wave called the Bow wave.

Particles trapped in the magnetic field may interact with molecules In the atmosphere to produce Aurora.

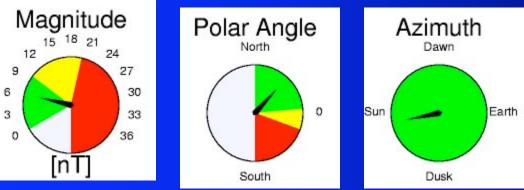


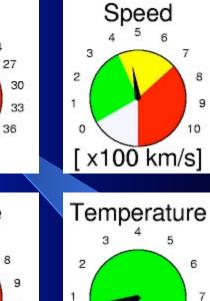
From <u>Stars</u>, James Kaler





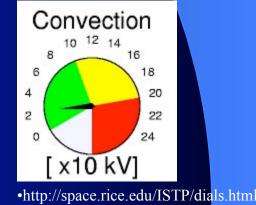
•Interplanetary Magnetic Field





•Voltage Across Polar Cap

[x100,000 K]



Interactions between the earth's magnetic field and a plasma can have spectacular results

• The northern lights (aurora borealis)



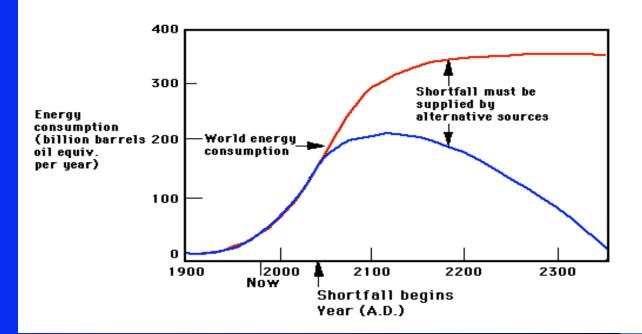
Photo by David Fritz http://dac3.pfrr.alaska.edu:80/~pfrr/AURORA/INDEX.HTM

More on Fusion Energy

- The goal of achieving controlled fusion energy has prompted much study into plasma physics.
- Fusion energy is a form of nuclear energy which is emitted when two light nuclei combine to form a single more stable nuclei.
- The sun and stars derive their energy from fusion.

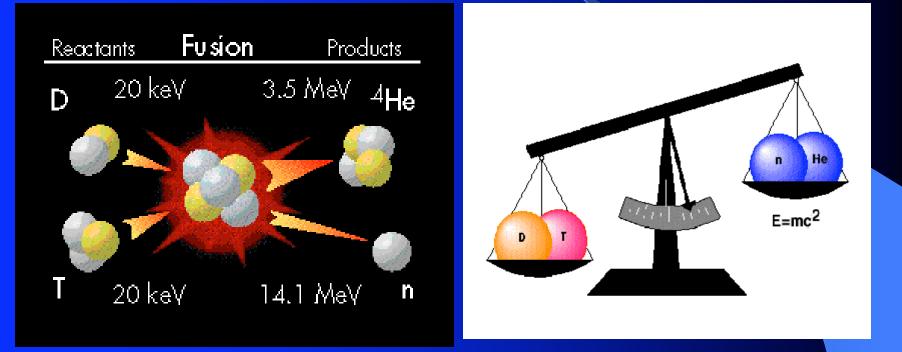
Why do we need new sources of energy?

population As the increases, so does the demand for energy. However, we have limited resources in the form of fossil fuels and this energy source must supplemented by alternative long term such sources, as fusion.



http://fusioned.gat.com/

Mass goes into energy in fusion reaction



The small reduction in mass of the products compared to the mass of the reactants leads to very large energy production.

Why is Fusion power needed?

Country	Consumption (kW-h/capita)
US	12000
Developed World Avg.	6000
World Avg.	1500
China	500
India	250

1990 Energy use per capita

•Projected change in consumption by increasing to world average



For more information see:

http://wwwofe.er.doe.gov/More_HTML/Artsimovich/PKKawPaper.html

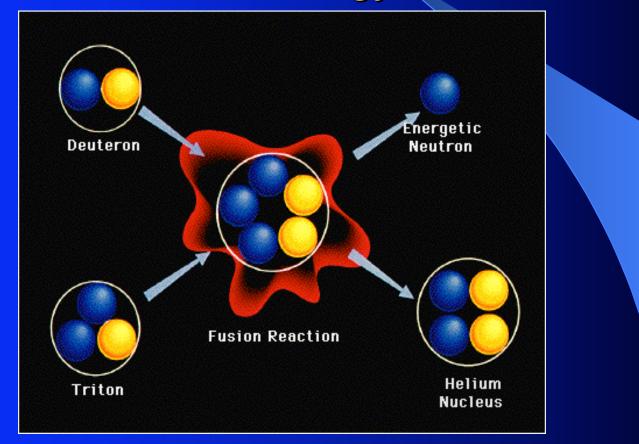
Fuel and waste products

• Fuel and waste for coal plants(most readily available energy source) vs D-T fusion plant

DAILY FUEL CONSUMPTION DAILY WASTE PRODUCTION 1,000 MEGAWATTS		
	COAL PLANT	D-T FUSION PLANT
F U E L	9,000 T. COAL	1.0 LB D ₂ 3.0 LB Li ⁶ (1.5 LB T ₂)
WASTE	30,000 T. CO ₂ 600 T. SO ₂ 80 T. NO ₂	4.0 LB He ⁴

http://www.pppl.gov

Under the right conditions, deuterium and tritium combine to form helium, a neutron, and fusion energy

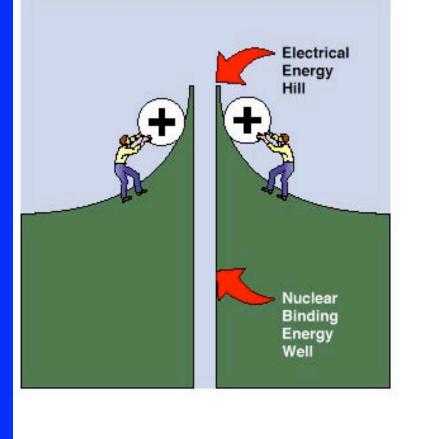


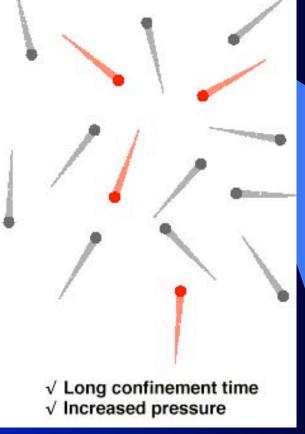
http://FusEdWeb.pppl.gov/

High temperatures and densities are needed

High temperature

Will they meet?

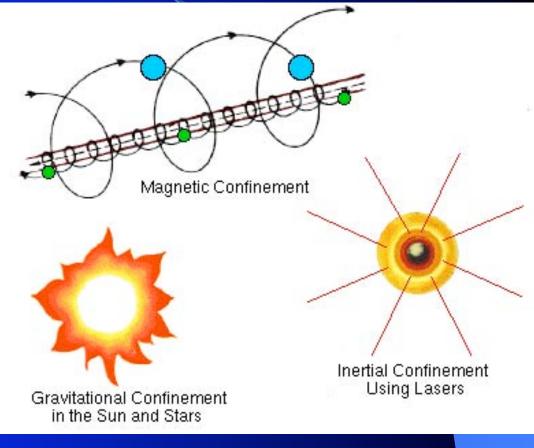




http://lasers.llnl.gov/lasers/education/talk.html

Methods for confinement

- Hot plasmas are confined with gravitational fields in stars.
- In fusion energy experiments magnetic fields are used to confine hot plasma, and inertial confinement uses lasers.



http://FusEdWeb.pppl.gov/

What must be achieved to obtain fusion energy?

- Contain a high temperature, T, high density, n, plasma for a long enough time, T, to achieve ignition (power out >> power in).
- A measure of plasma performance is thus given by:

nTτ

density * temperature * confinement time

Two major approaches to fusion (D–T)

Magnetic confinement

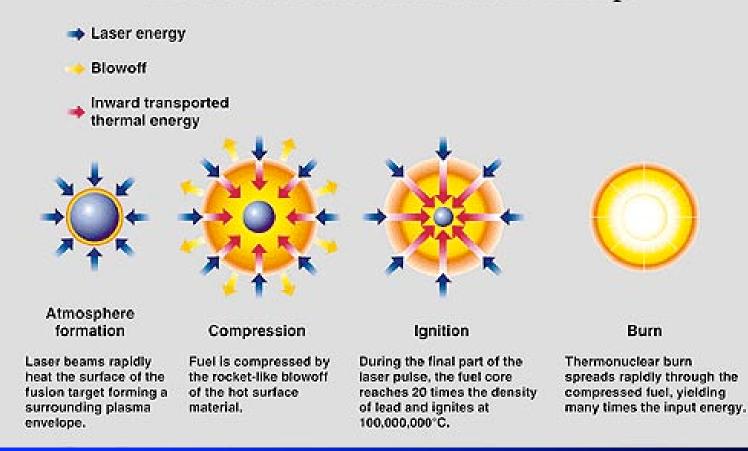
- Temperature ≈ 10⁸ °C (10 keV)
- $n\tau\approx 10^{15}$ atoms $\cdot seconds$ / cm 3
- $\tau \approx 10$ seconds (magnetic "bottle")
- $n \approx 10^{-14}$ atoms / cm⁻³ (10⁻⁵ times the density of air)

Inertial confinement

- Temperature $\approx 10^{8} \circ C (10 \text{ keV})$
- $n\tau \approx 10^{-15}$ Atoms ·seconds / cm ³
- $\tau \approx 3 \times 10^{-11}$ seconds (microexplosion, inertial "bottle")
- $n \approx 3 \times 10^{25}$ Atoms / cm ³ (12 times the density of lead!
- ~ 1000 times the density of liquid DT!)

Controlling Fusion using Inertia

The Inertial Confinement Fusion Concept



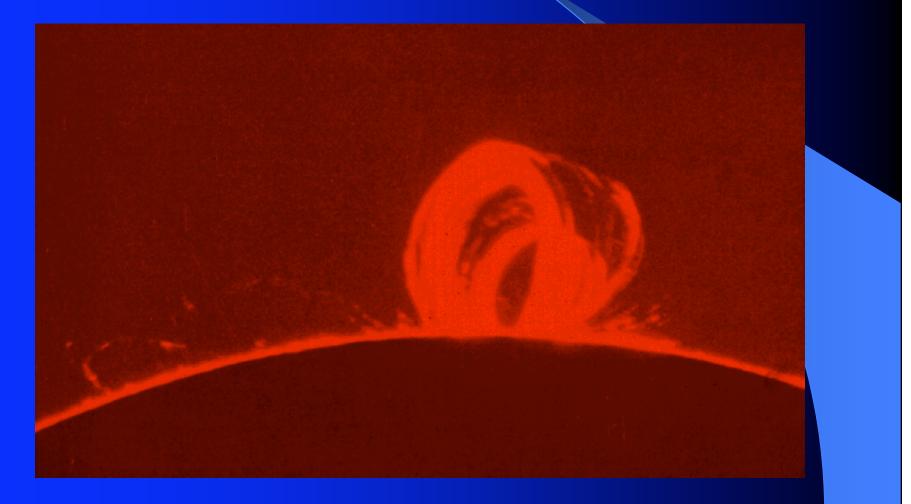
http://www-lasers.llnl.gov/lasers/nif/nif_ife.html#fusion

Direct vs Indirect Drive



Particles in a Magnetic field

Moving charged particles will feel a force when placed in a magnetic field. $\mathbf{F} = \mathbf{q}\mathbf{v} \times \mathbf{B}$.

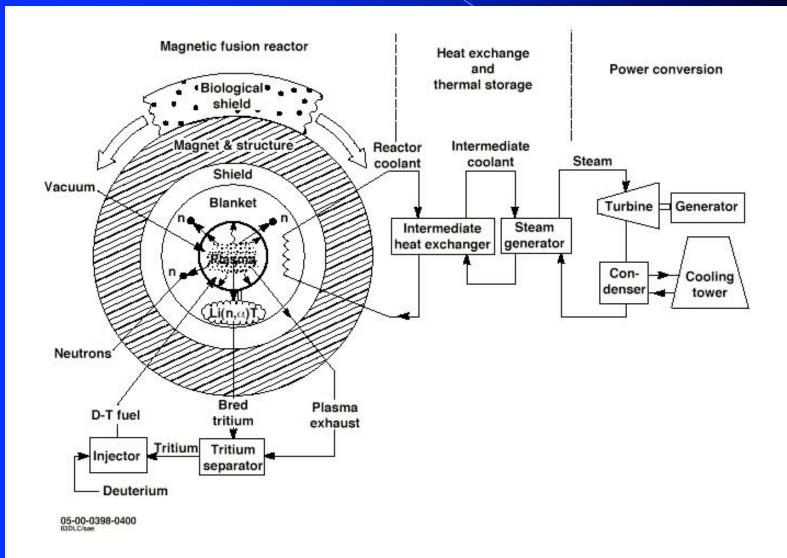


Much research has gone into controlling fusion with magnetic fields

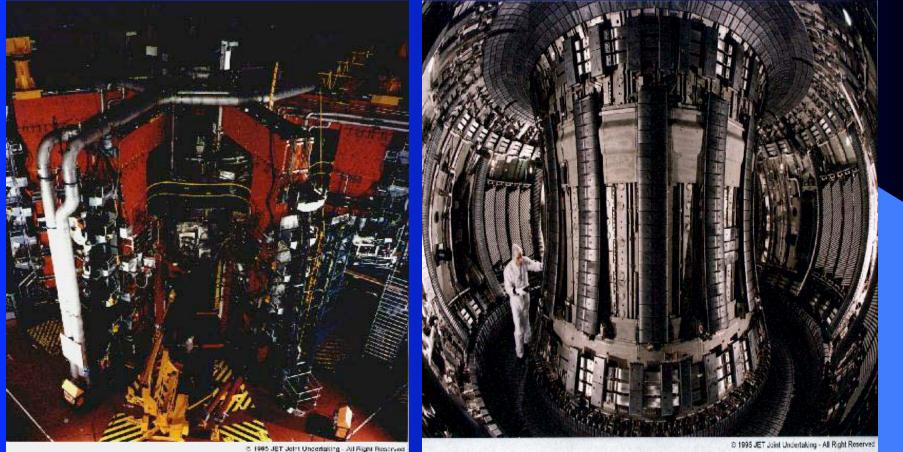
- Most magnetic confinement devices in use today have a toroidal shape.
- Large magnetic fields are created by driving currents through coils wrapped around the torus.



A simplified power plant schematic shows the reaction chamber, magnets and confining structure, and the proposed lithium blanket that would be used as a heat transfer fluid



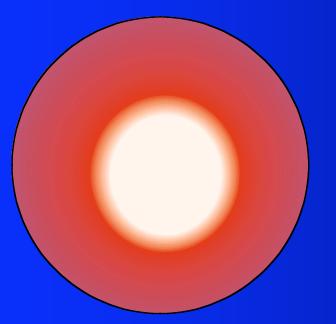
Joint European Torus: the largest confinement device ever built

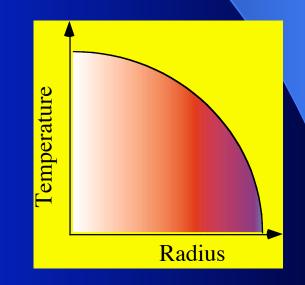


http://www.jet.uk/

Need to control temperature and density

• We need the core hot enough for fusion, yet the edge cool enough not to melt the walls





But nature abhors gradients:

- Whenever a slope (gradient) gets too steep, nature finds a way to flatten it out
 - Mountains get eroded
 - sand and snow avalanche
 - turbulence grows to flatten steep slopes in plasmas
- We need to control the turbulence

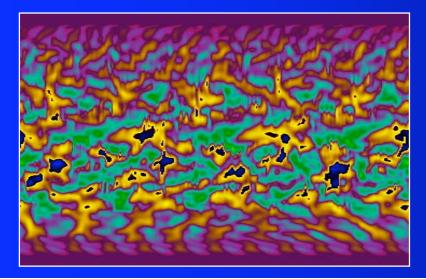
Turbulence moves things down the slope

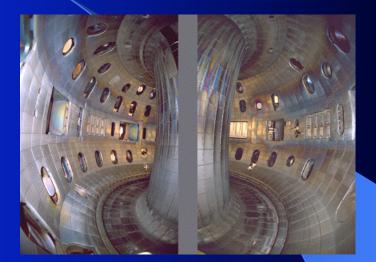
 The turbulence swirls (eddies) move the heat and density toward the edge



Challenges on the path to Fusion

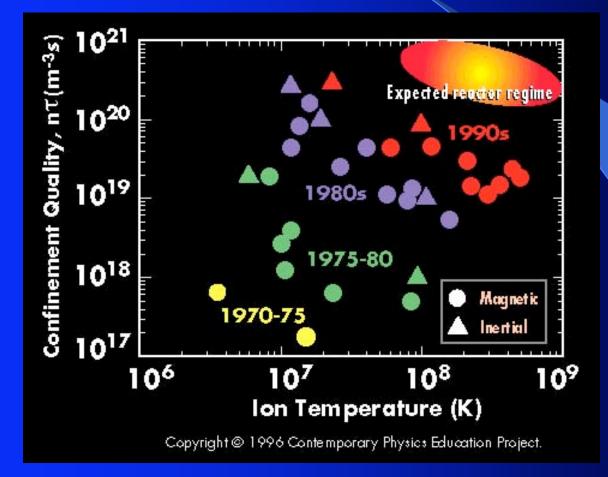
- Heating
- Fueling
- Confinement
 - Plasma physics is on the leading edge of technology







Progress towards fusion energy



http://FusEDWeb.pppl.gov/CPEP/Chart_pages/6.Results.html

Web References

Fusion energy and plasma educational sites

- http://FusionEd.gat.com/ General Atomics
- http://FusEdWeb.pppl.gov/ Princeton Plasma Physics Laboratory
- http://lasers.llnl.gov/lasers/education/ed.html Lawrence Livermore National Laboratory
- http://www.jet.uk/ Joint European Torus
- http://www.ornl.gov/fed/fedhome.html/ Oak Ridge National Lab
- http://www.ornl.gov/fed/theory/Theory_Home_page.html
- http://www.ornl.gov/fed/mhd/mhd.html/ Oak Ridge National Lab

• Astrophysics sites

- http://umbra.nascom.nasa.gov/spd/ NASA Space Science
- http://www.seds.org/billa/tnp/ The Nine Planets
- http://www.stsci.edu:80/ Space Telescope Science Institute
- http://bang.lanl.gov/solarsys/ Views of the Solar System
- http://www.gi.alaska.edu/ Geophysical Institute (Aurora and Sprite info)
- http://www.sec.noaa.gov/ NOAA Space weather site

Plasma is Coming to Denver Fall of 2005

- APS/Division of Plasma Physics meeting
- Teachers Day
- Student
- Plasma Expo (interactive booths)
- See http://fusioned.gat.com -bookmark
 Keep an eye on (Upcoming events)